

# $\text{Ly}\alpha$ emission as a sensitive probe of feedback-regulated LyC escape from dwarf galaxies

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1

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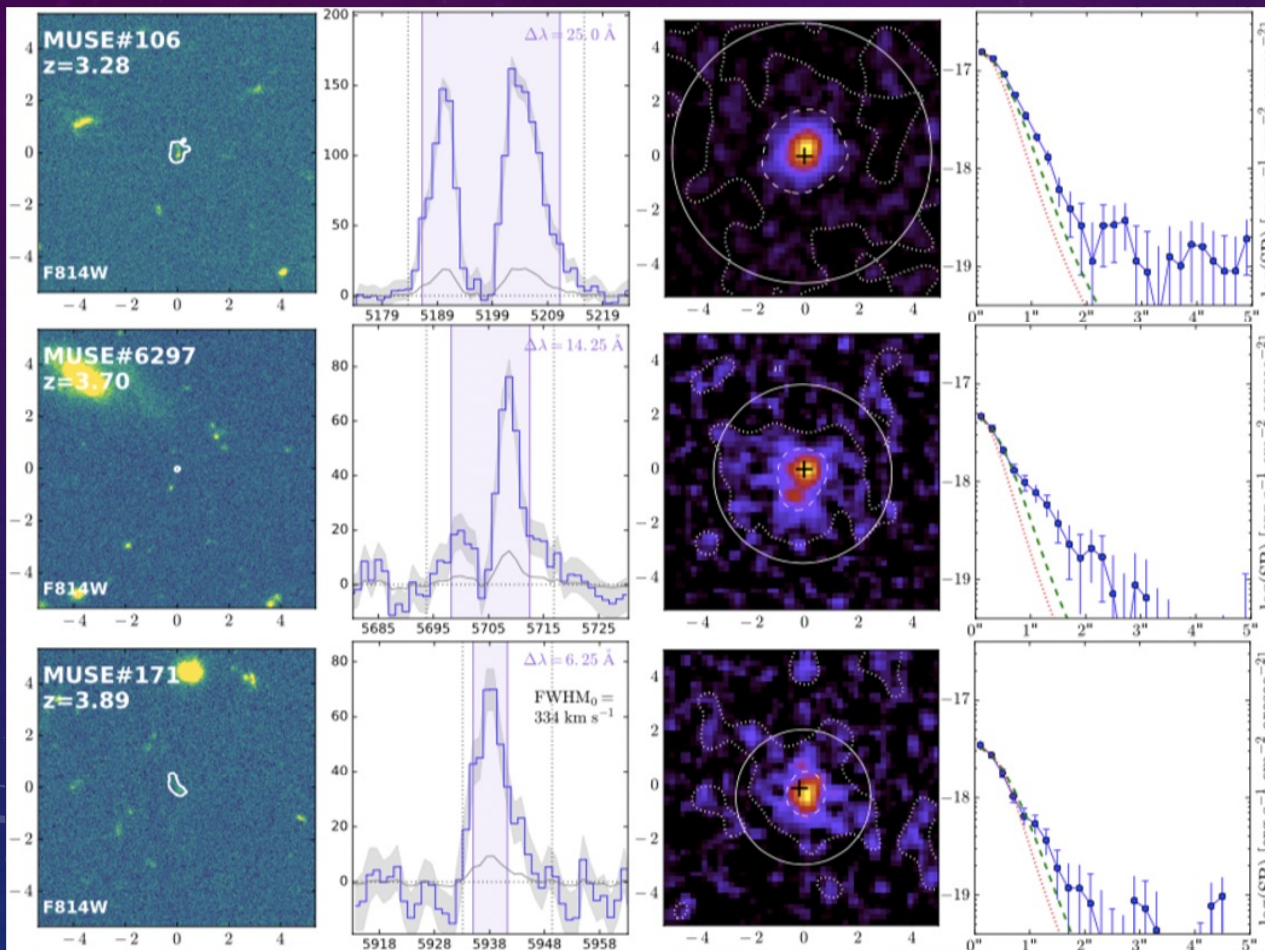
# MUSE UDF SURVEY OF $\text{Ly}\alpha$ HALO

UV images

$\text{Ly}\alpha$  spectra

$\text{Ly}\alpha$  images

SB profiles



- $\text{Ly}\alpha$ : tracer of star-forming galaxies
- Recombination emission
- Resonant scattering
- $\text{Ly}\alpha$  halo (LAH) is more extended than UV halo.
- Large variations of  $\text{Ly}\alpha$  signature between objects.



# PANDORA Martin-Alvarez+23

- Cosmological zoom-in simulation suite of dwarf galaxies
- RAMSES-RT
- Supernovae (SN) feedback, radiative transfer (RT); magneto-hydrodynamics (MHD); cosmic ray (CR)

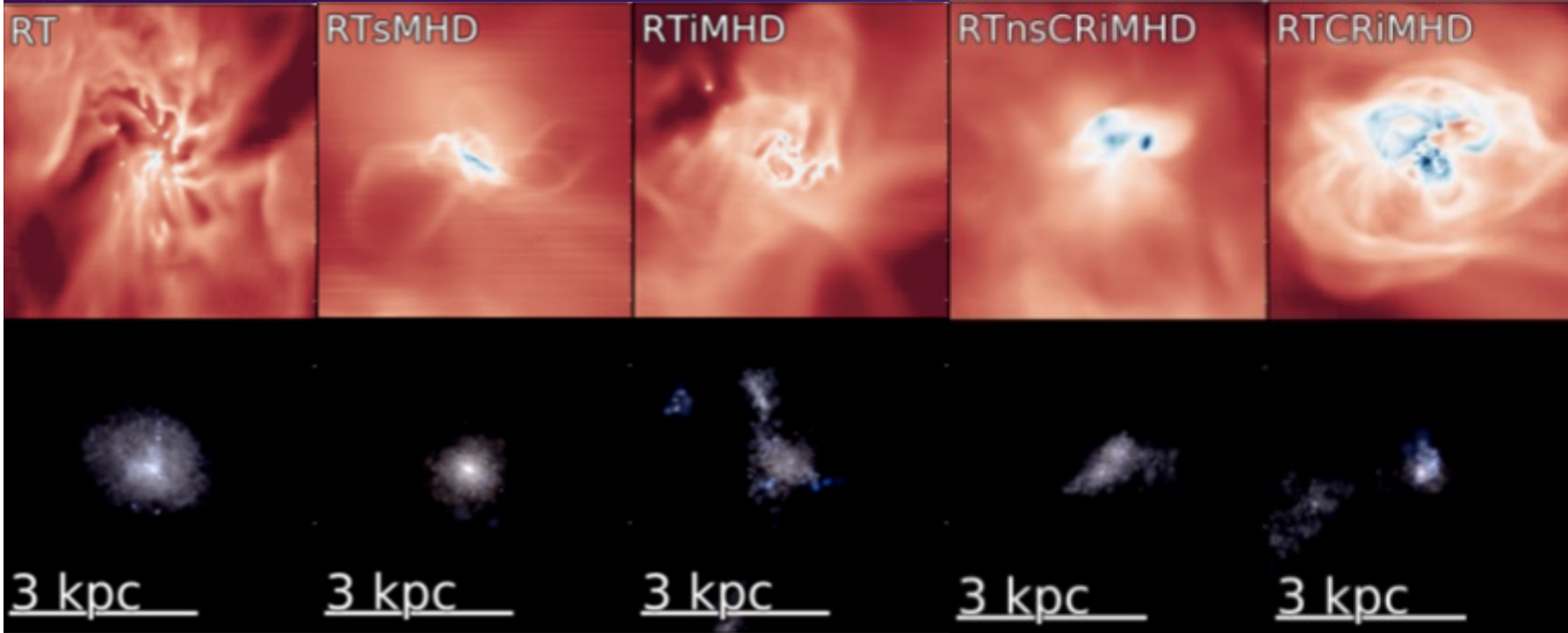
RT

RTsMHD

RTiMHD

RTnsCRiMHD

RTCRiMHD



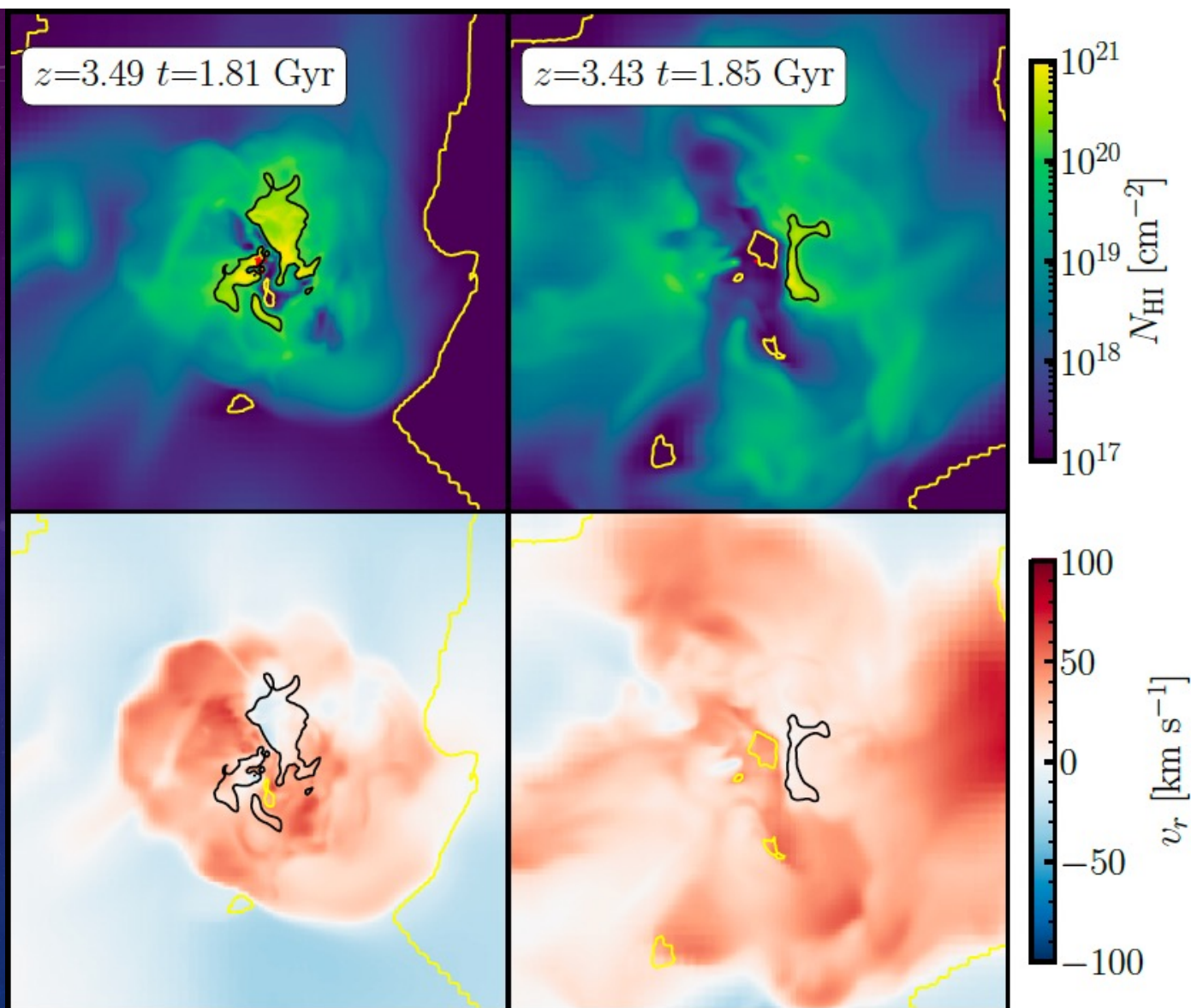
# LY $\alpha$ POST-PROCESSING



- Postprocessing with **RASCAS** (<http://rascas.univ-lyon1.fr/>)
- Sample with 200,000 photon packets.
- Resonant scattering by HI.
- Photons
  - escape from the galaxy
  - get absorbed by dust
- Synthetic observations along 108 LOSs

# GAS PROPERTIES

- RTCRiMHD simulation
- Expansion of outflow





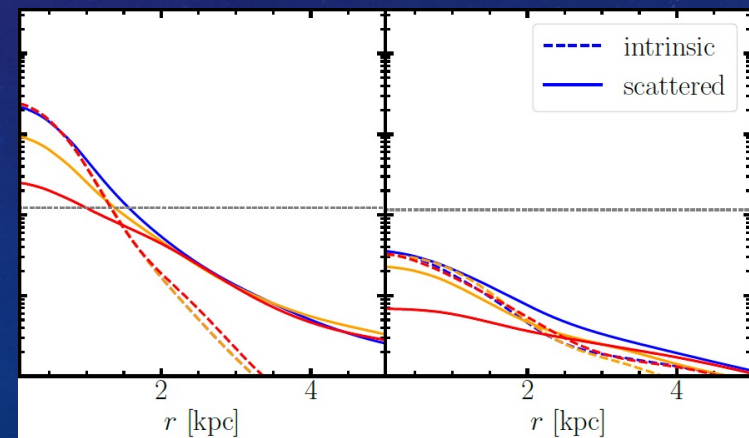
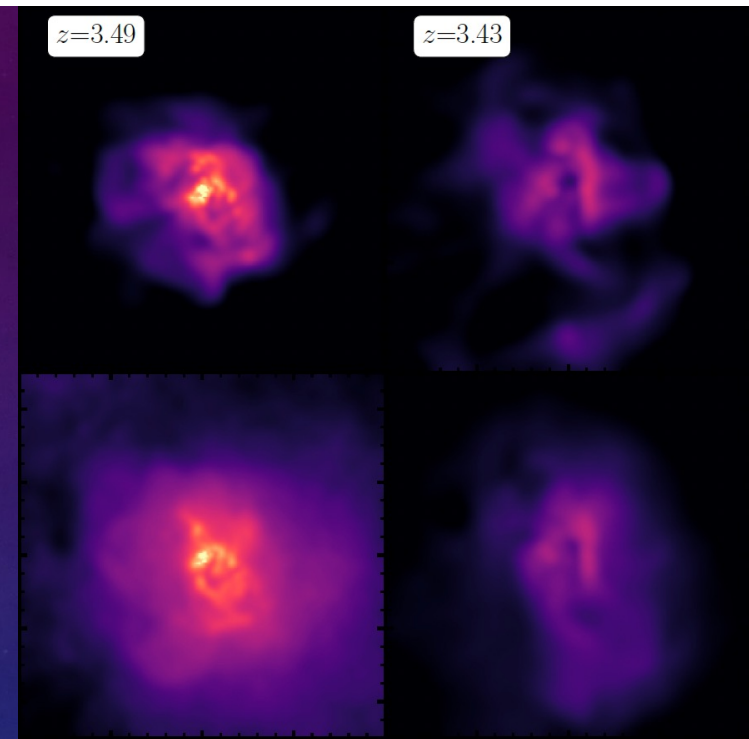
# LY $\alpha$ IMAGES

- Scattered Ly $\alpha$  images are more extended than intrinsic Ly $\alpha$  ones
- Along different LOS:
  - Similar intrinsic SB profile
  - Different scattered SB profile
- Weaker broadening effect when the wind is fully developed

Intrinsic  
Ly $\alpha$

Scattered  
Ly $\alpha$

SB

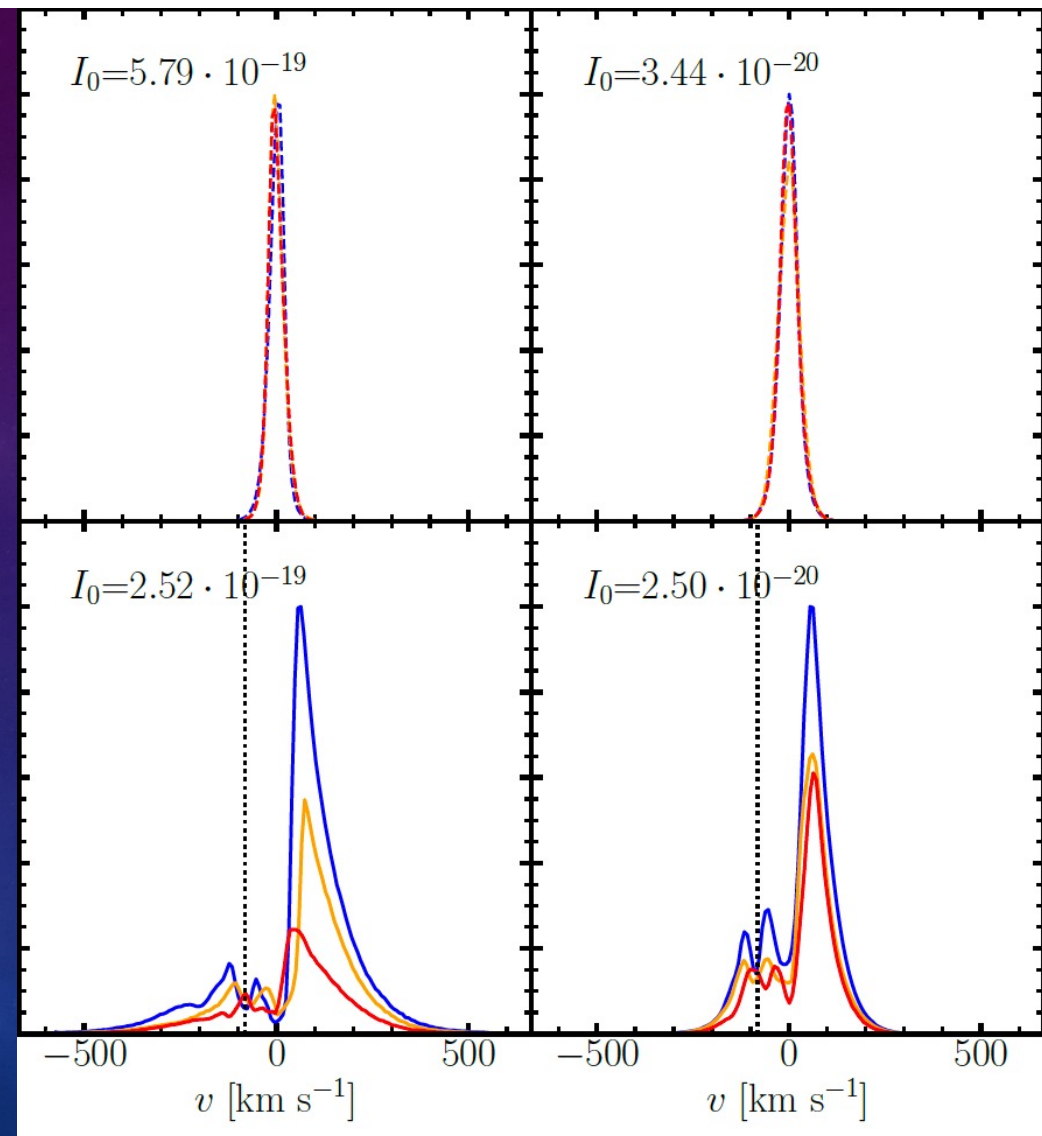


# LY $\alpha$ SPECTRA

- Double (triple) peaked profile
- Prominent red peaks -> outflow
- $v_{sep}$ : velocity separation between red and blue peaks
- R2B: size ratio of the red to blue peak

Intr Ly $\alpha$

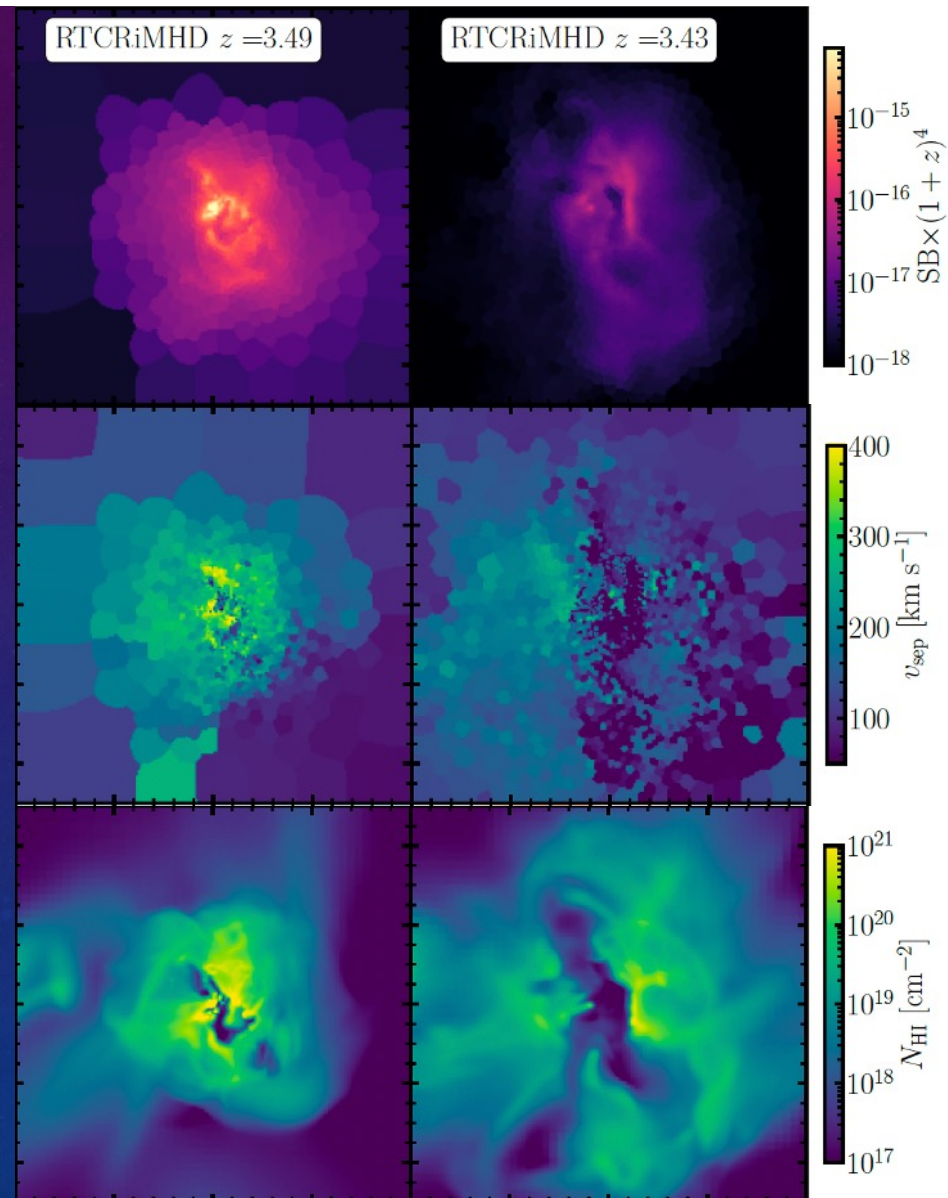
Scat Ly $\alpha$





# Inferring the gas distribution from Ly $\alpha$

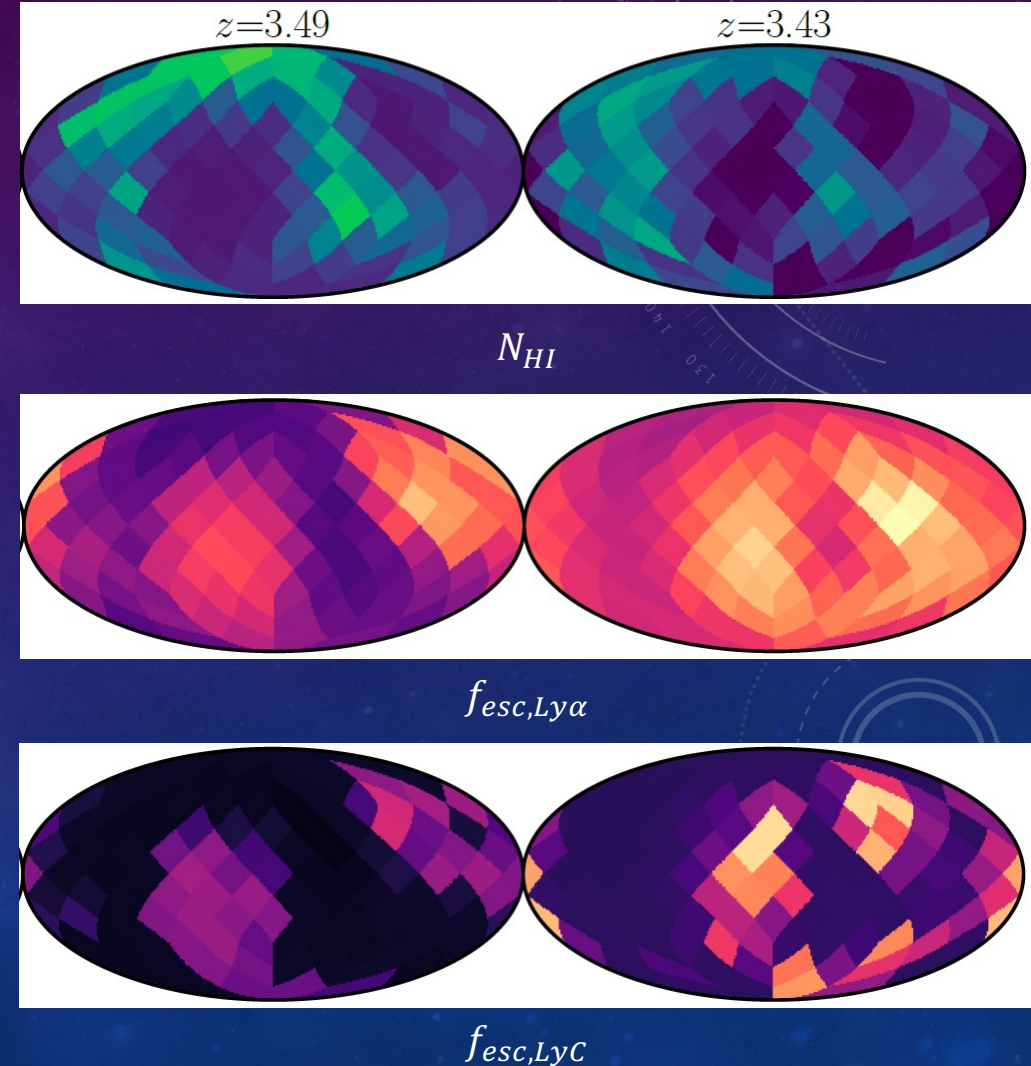
- Spatially resolved Ly $\alpha$  profile
- Voronoi tessellation
- Calculate SB,  $v_{sep}$  for the spectrum in each Voronoi bin
- $v_{sep} \rightarrow N_{HI}$





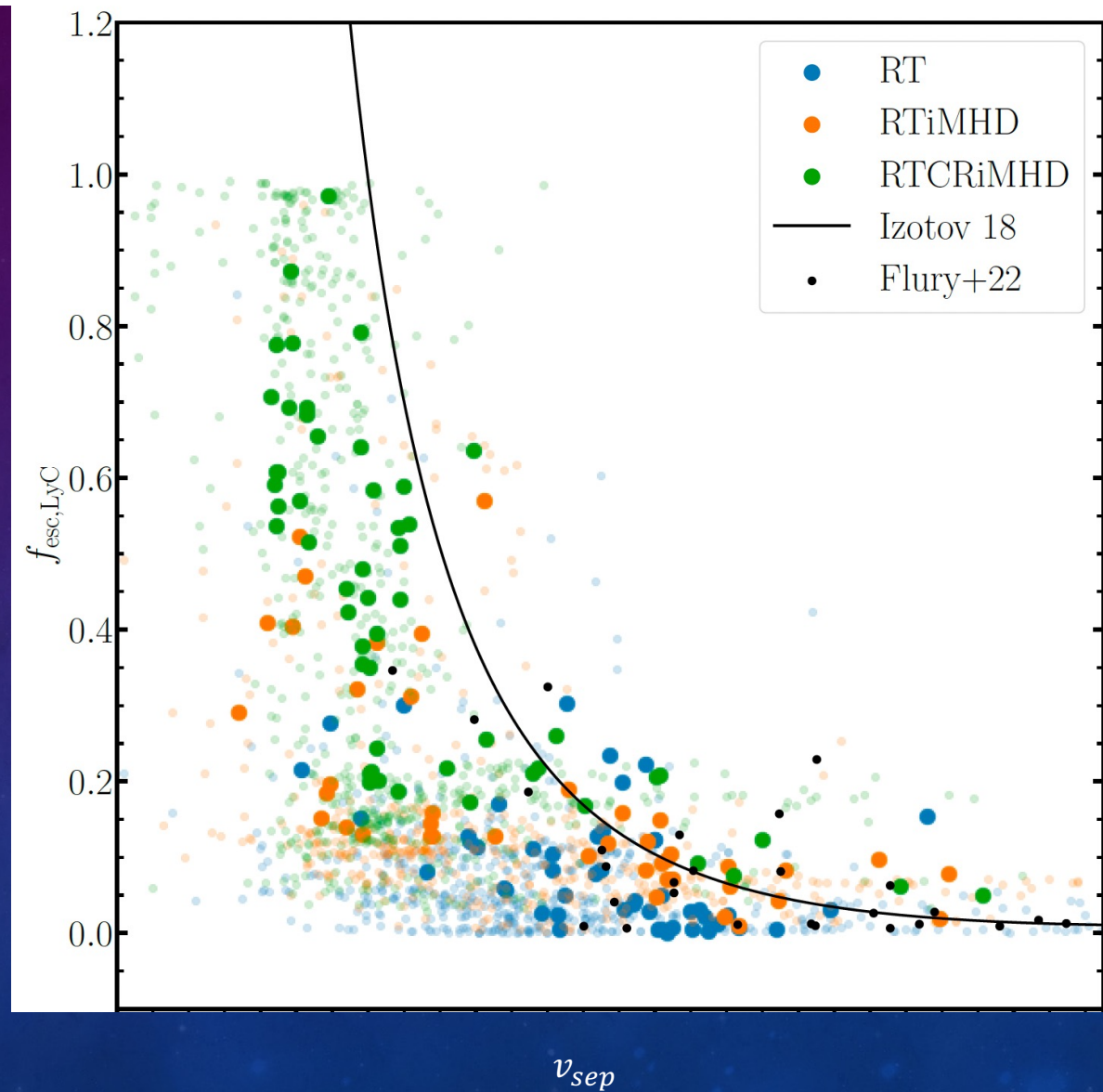
# Anisotropic Ly $\alpha$ and LyC escape

- Feedback  $\rightarrow$  low HI column density channel  $\rightarrow$  Ly $\alpha$  and LyC escape
- The distribution of  $f_{esc,LyC}$  has a larger contrast compared to  $f_{esc,Ly\alpha}$
- Different escape mechanisms



# Decipher LyC escape with all physical models

- 3 simulations span an universal relation
- $v_{sep} - f_{esc, LyC}$ : reproduce the Izotov18 relation.





# Conclusions

- Dwarf undergoing starburst show extended Ly $\alpha$  emission and red-peak-dominated spectra.
- Spatially resolved Ly $\alpha$  opens a new window to investigate the neutral gas kinematics.
- Ly $\alpha$  and LyC escape differently and depends strongly on the LOS.
- Different simulation runs have diverse Ly $\alpha$  and LyC features, but they all follow universal relations on the parameter space.